Separating detection and catalog production,

a step towards Algorithms for Large Databases and Vice-versa.

Mohammad Akhlaghi

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October 19th, 2016. 26th ADASS. Trieste, Italy.

A vote for the key topic





Please participate in the selection of key themes for ADASS 2016 by proposing an...

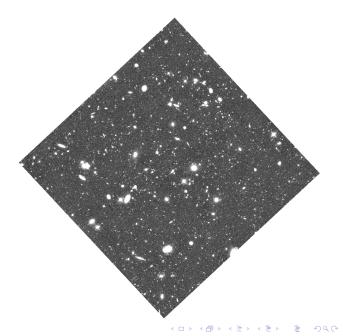
Votes	Answer	Results	
1 O Mohammad Akhlaghi Reduction/analysis algorithms fo comming era of massive surveys database designers and the algor the decisive factor in scientific pr users/scientists to be more creat algorithms can greatly enhace sc	r large databases and (e.g. LSST, SKA), the ro rogress. Systems that ive with the reduction	94.8% vice versa: In the ole of the adopt becomes allow/encourage	2 Manag this top project do we can we marchi

https://app.appgree.com/#!talk/927227906403405825/1659496956281499649

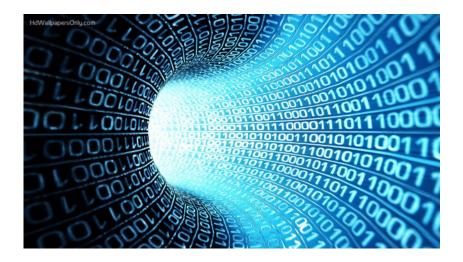
Raw data is not directly usable for analysis.

Each data element (pixel) only has a position and a value.

Example image: HST UDF field in F775W (Illingworth+2013).



So, we reduce the low-level data ...



... into high-level constructs

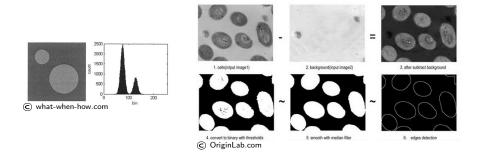
Each row is one identified subject (group of pixels).

Each column is one derived property (e.g., position, magnitude, or ellipticity).

6473 53.12683058 -27.79528847 9037.4358 4796.1568 99.881 99.006 27.284 27.258 27.316 26.913 99.885 26.233 0.000 8.842 0.03 29.437 28.377 27.79547832 4959.4948 30.092 29,169 29.285 28.128 29.951 4763.9618 -27.79545453 5272.0120 4765.9610 -99,000 99,000 99.000 29.786 29.453 29.164 28,989 28.533 28.343 8,889 0.000 0.000 8,453 9.201 0.183 -27.79548546 9319.3198 4762 9240 -99.888 28.368 28.232 99.000 28.252 0.000 0 000 0.049 0.055 24.343 -27.78565732 6224.4228 29.358 28.268 79548315 6611 1448 6486 9766 53.15723140 27.75778472 5888.9988 9296.1028 99.000 29.995 29.866 28.211 28.448 28.235 99.888 758 38.882 1.688 0.844 0.047 6481 6366 -27.78616468 4268 6448 5880 3788 99.000 29.347 29.482 28.188 28.364 27.882 8.888 0.000 0.196 -99.881 6482 53.13494109 79548139 8176.4728 4763.3600 -99.000 99.88 99.000 38.464 29.448 29.342 29.893 29.166 99.000 99.886 99.000 8.899 0.000 0.000 8.289 0.089 0.090 6483 78628184 6977.5588 5876.5188 -99.000 99.881 99.000 38.151 27.805 27.768 27.739 27.782 8.000 0.000 0.000 8.487 0.046 4588 53.15318829 79347772 6239.3968 5003.4540 27.967 27.865 26.583 26.342 26.20 26.100 26.053 0.370 0.039 0.022 6485 53.15466929 79548426 6082.2418 4762.6298 -99,000 -99.881 99.000 .311 31.146 31.598 30.788 30.011 38.154 30.289 8,000 0.000 0.00 8.256 0.185 6486 28.415 28.242 29.544 29,491 8.846 0.039 31.089 31.058 .78574961 8134.0888 5931.1590 30.446 30.915 1.032 8.288 0.173 0.245 70506074 7709 1710 29,467 0.000 \$928.9400 29.324 53.16847727 24.077 -27.78629732 5465.3628 26.125 24.756 24.188 .78626722 5998.2218 29.958 28.743 0.658 0.059 53 12981124 -27.79553694 8805.9588 4756 7386 .00 hht -99.881 99 006 28.533 27.989 28.76 99.888 27 996 8 888 0.000 0.00 8.661 0.058 27 79556488 8794 1748 4753 4738 99 666 28 724 8 886 79553126 7958 8991 29.159 53.13728748 .79552769 7927.3938 4757.7668 99.000 99.88 99.000 38.660 29.58 99.000 99.888 99.000 8.888 0.000 0.000 0.156 649/ 53.14388811 79552547 7328,1258 4757.9466 .99.000 99.88 99.000 31.285 10.784 30.468 31.016 29.685 30.336 99.000 8.889 0.000 0.000 8.481 0.186 8.169 .79553444 4208.3400 4755.9860 -99.000 -99.881 99.000 29.988 29.859 29.785 29.895 8.999 0.000 53,18153224 78565751 3230,0631 99.000 99.000 99.000 28.179 27.987 28.006 28,061 27.493 27.478 27.150 29.85 0.145 0.083 0.10 6586 78571246 6191.6458 5935.2730 99.000 99.888 29.311 28.745 27.988 27.884 28.861 27.130 27.025 38.004 0.890 0.843 0.048 78559876 3868.6838 5949.0878 99.000 28.962 99.000 29.598 29.286 28.838 28.536 27.858 27.868 38.293 0.885 30.878 0.135 0.10 99.000 29.954 29.962 29.375 99.888 0.000 0.000 0.178 0.189 27 70555016 7000 1070 99,000 28.718 28.671 8,882 0.842 0.000 0.043 27.561 0.230 27 78635410 6002 7288 6588 -27.78633660 5344.5898 5850 1368 99 666 29.188 27.741 27.545 27.451 0.000 8.616 27 79556925 2682 3628 4750 7778 .99 .000 99.88 99.666 99.886 29.648 29.582 29.663 29.127 99.886 99.000 8.88 0.000 0.004 31.922 0.128 53.14738483 27.79559643 6864.8978 4749 3488 99.666 38.474 8.889 8.886 -99.881 0.000 78585466 8139,688 5918.5538 .99.000 -99.88 99.000 10.411 30.34 38.584 30.178 38.178 38.211 30.368 8.888 0.000 0.000 8.341 8.19 53.17808141 78478449 3596.3698 6945.7198 99.000 99.888 99.000 28.666 28.482 28.462 28.149 38.355 30.395 30.902 8.161 0.08 6512 4105 53,19692300 79559973 1596,8338 4746.8360 - 99,000 99.001 99.000 29.907 99.000 99.000 99.00 99.000 8.000 0.000 0.000 0.094 6514 6403 .78586195 6809.7838 5917.4750 -99,000 99.881 99.000 38.685 29.486 29.024 29,491 29.284 29.228 28.845 28.759 8.000 0.000 0.000 8.470 .79563710 8088 1948 99.000 8.291 53.19144757 27.79561505 2178.0820 99.881 99.000 29.848 30.296 38.159 28.981 28.325 99.000 0.000 0.000 0.145 0.18 53.14929275 -27.78611953 6652.7418 26.837 29.954 0.887 31.599 31.962 -27.78619325 2816.1388 30.141 29.043 99.000 0.000 .27 78523322 7911 2698 5993 1000 99 005 99 000 27 501 26.516 26.348 8 848 0.016 -27 79565626 4878 6248 4741 3118 -99 006 28 489 8 348 .99 888 .27 78596998 5863 8578 5984 2948 99.888 27 499 -27.78598786 6953.9928 5911.9948 30.694 38.843 .99.000 -99.888 .99.666 0.000

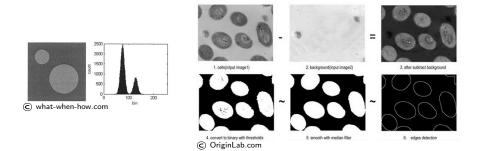
(UVUDF catalog, Rafelski et al. 2015)

When there are sharp and high S/N edges, a sufficiently high threshold can avoid the noise \rightarrow Signal-based detection.



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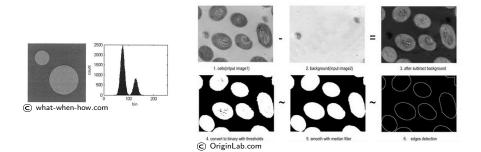


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Outline is:



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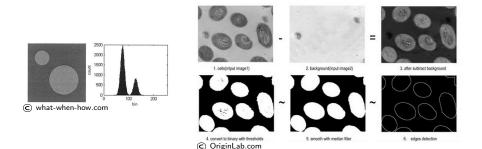


Outline is:



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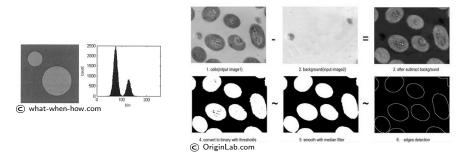
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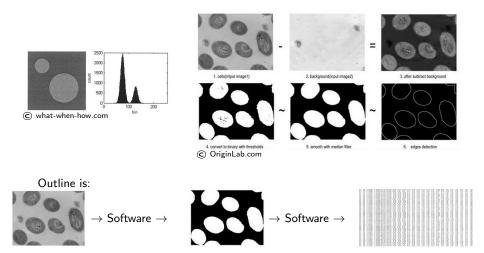
Outline is:



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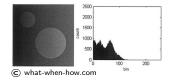
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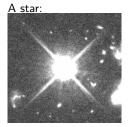


But astronomical objects ...

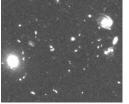
... don't have such sharp edges.

... can have a huge diversity of shapes and sizes.



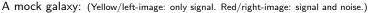


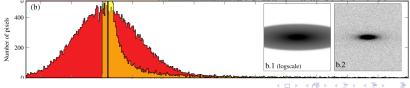
Some galaxies:



A main-belt comet:



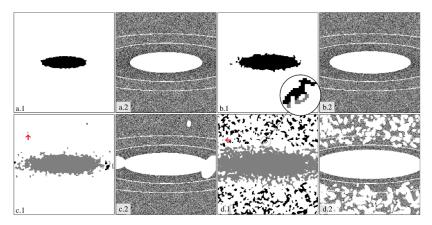




SAC

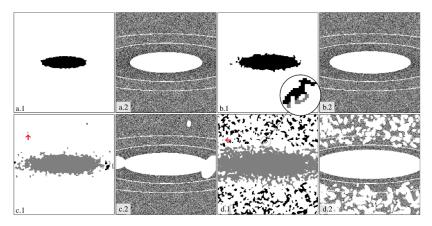
So for astronomical objects ...

... a threshold designed to avoid the noise (signal-based detection) will miss a lot of the signal. Decreasing the threshold will result in many false detections. So our only hope is to try modelling the brighter parts.



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Detection and catalog production heavily intertwined through the parametric models.

Outline of most common software

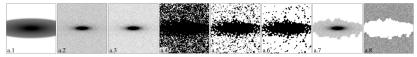
So this is the most common procedure:



As a result:

- Catalog production is computationally expensive.
- Decreases modularity, or creativity.

In Sydney (25th ADASS), we introduced a new noise-based detection. The threshold is below the Sky value ($\sim -0.5\sigma$, not designed to avoid the noise). Noise is separated from signal (detection) by exploiting the signal's contiguity.

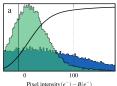


The threshold is the solid vertical line:

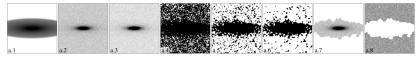
(In this mock image, the Sky value is on 0)

See:

Akhlaghi and Ichikawa (2015, ApJS 220, 1. arXiv: 1505.01664).



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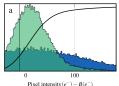


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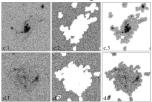
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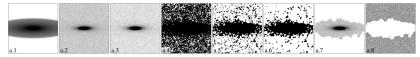
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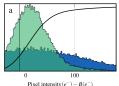


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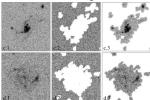
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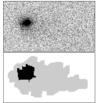
Akhlaghi and Ichikawa (2015, ApJS 220, 1. arXiv: 1505.01664).



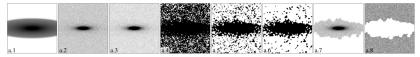




A comet (Hsieh+2013):



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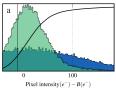


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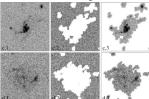
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See:

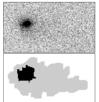
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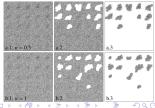
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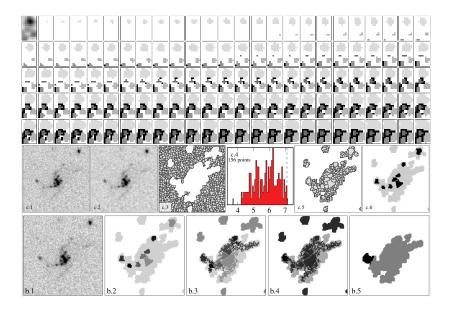
A comet (Hsieh+2013):



Mock faint/small galaxies



Finding true peaks and segmentation



In Gnuastro, NoiseChisel is only in charge of the detection (creating a labeled image), and MakeCatalog is in charge of creating a catalog from that input.



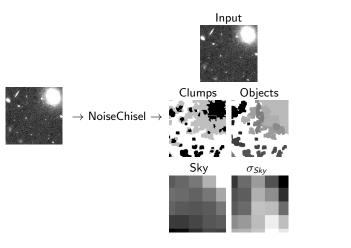


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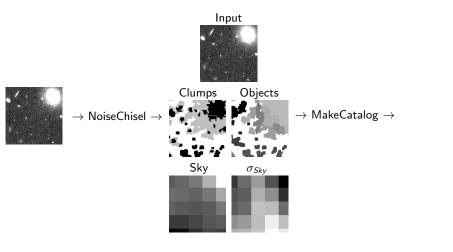




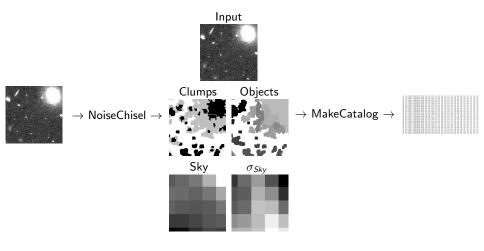
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Basic idea:

Each image (input, Clumps, Objects, Sky, and σ_{sky}) is a separate input into MakeCatalog. (The Clumps are optional)

Therefore

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The modularity (separating detection from catalog production) greatly simplifies pipeline development.

And, most importantly ...

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Therefore

The modularity (separating detection from catalog production) greatly simplifies pipeline development.

And, most importantly ...

empowers the users in using the data-base as they please.

Initial objects segmentation map ...



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Initial objects segmentation map ...



... can be broken up into individual segmentation maps for each label ...

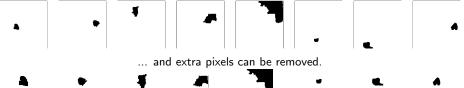


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Initial objects segmentation map ...



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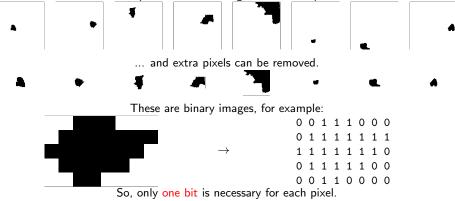


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Initial objects segmentation map ...



... can be broken up into individual segmentation maps for each label ...



Deblending and matching

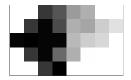
Deblending (when necessary) can be defined as 1-byte (256 layers) or 2-byte (65536 layers) integers.

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0	0	200	170	100	80	0	0
0	200	250	180	80	90	50	30
250	255	250	250	70	50	40	0
0	200	250	150	80	1	0	0
0	0	200	255	0	0	0	0

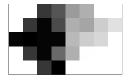


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0	0	200	255	0	0	0	0

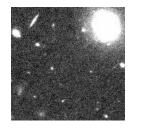


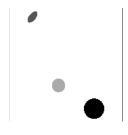
In a similar way, the pixel maps can be warped and/or convolved to match images with other pixel or spatial resolutions (from other surveys).

Aperture photometry

Aperture photometry only needs detection for the Sky and σ_{Sky} .

In Gnuastro, MakeProfiles is in charge of building profiles (apertures in this case) on an image.



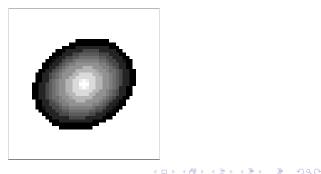


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Curve of growth

Curves of growth segmentation maps (elliptical annuli) can also be created easily with MakeProfiles.

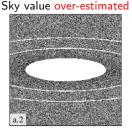
MakeProfiles input (to define each annulus): 1 22 22 5 1 0 30 0.8 1 1 2 22 22 5 2 0 30 0.8 2 1 3 22 22 5 3 0 30 0.8 3 1 (Columns are: ID, X, Y, function, radius, func. param, PA, axis ratio, value, truncation)



High completeness catalog in multi-band data

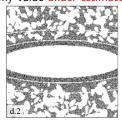
Completeness and purity (no contamination) are anti-correlated. For the image below, assume this is a real image with faint galaxies in the background.

Threshold: 2σ High purity (no false detections) Low completeness (faint objects lost)



Threshold: 0.1σ

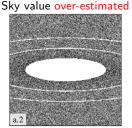
Low purity (many false detections) High completeness (faint objects detected) Sky value under-estimated



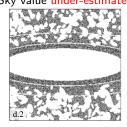
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Threshold: 2σ High purity (no false detections) Low completeness (faint objects lost)



Threshold: 0.1σ Low purity (many false detections) High completeness (faint objects detected) Sky value under-estimated



When the science needs multiple detections (like dropouts with detections in
redder filters), the low purity will be corrected and we end up with a high
completeness by taking this strategy:Reasonable completeness and purity:High completeness:
Sky and σ_{sky} .Sky and σ_{sky} .Segmentation map(s).

Even more:

Custom columns:

- Each pixel only has position (x and y) and value (v).
- So new columns can be defined at runtime.
- Or MakeCatalog can link with dynamically loaded libraries (plugins) to define more columns by the users.

Over-lays

The segmentation maps can be uploaded to online viewers like VisiOmatic (Poster P8.4, Bertin et al.) for users to visually check their objects online.

Different methods to use clumps and objects

NoiseChisel's segmentation is just one approach to using clumps. Other science cases can use them differently for other purposes.

It is not unique to Gnuastro

 This separation can greatly simplify any calculation, not just users of Gnuastro/MakeCatalog.

NEWS: GNU Astronomy Utilities (Gnuastro) 0.2 released

- Conforms with GNU Coding Standards.
- Now installs shared libraries (for C and C++).
 - Emphasis was mainly on robust build and documentation.
 - Most interesting functions are still locked within programs and will be liberated soon.
 - The quantity and quality of libraires will greatly improve with future releases.
 - Libraries will soon be changed to LGPL.
- Comes with 14 programs: Arithmetic, ConvertType, Convolve, CosmicCalculator, Header, ImageCrop, ImageStatistics, ImageWarp, MakeCatalog, MakeNoise, MakeProfiles, NoiseChisel, SubtractSky, Table.
- Complete documentation in various web, print and command-line formats.

Webpage: https://www.gnu.org/software/gnuastro/ Please come help us make it better for everyone.

