

How can we assess aperture photometry degradation after HiPS transformation?

Context:

- HiPS is an emerging IVOA standard designed to cope with pressing challenges for data visualisation and distribution
- Initially developed at CDS (figure 1), it is based on HealPix tessellation of the sky which defines cells of equal area on all sky
- Here we raise the question of potential quality degradation in producing the HiPS version

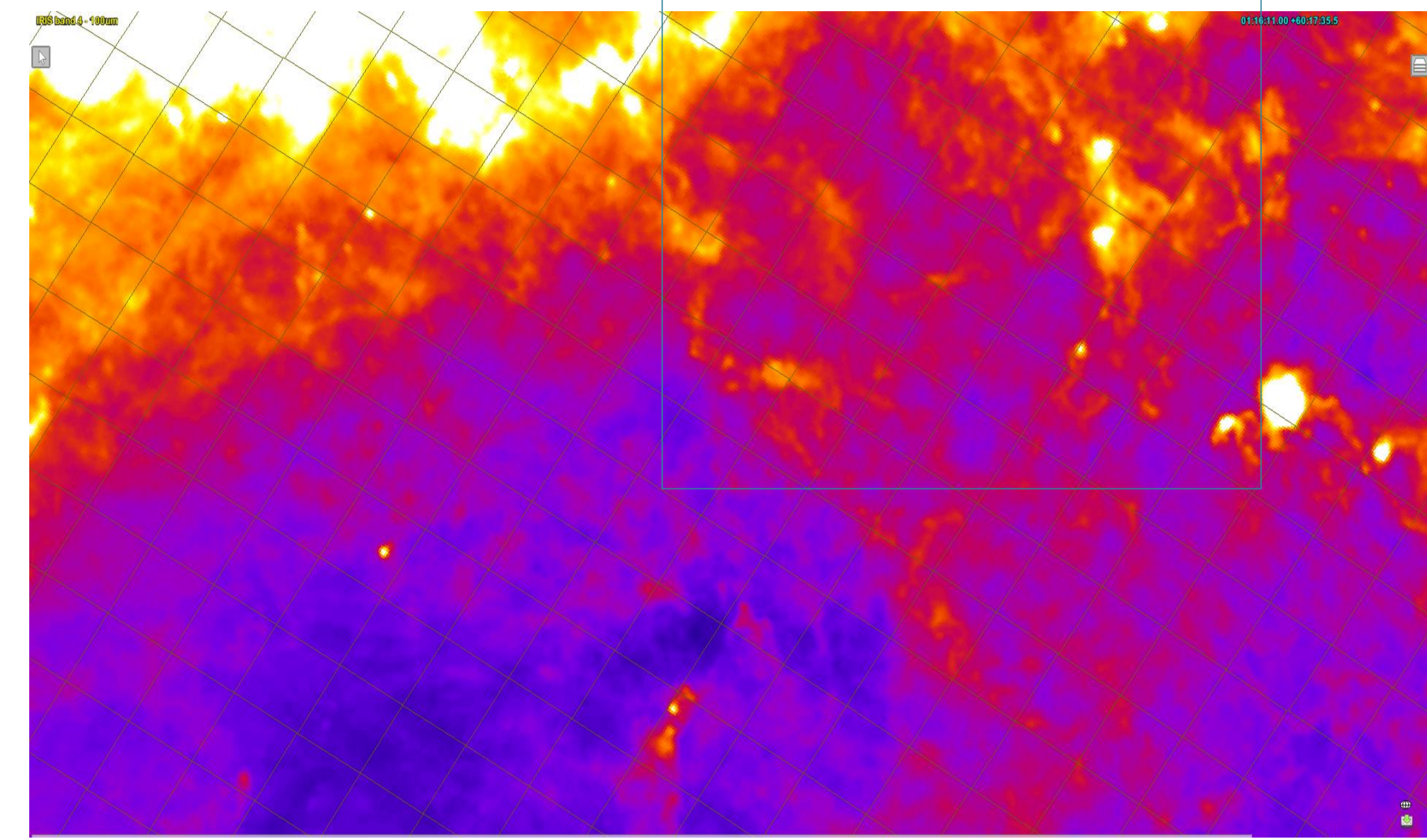
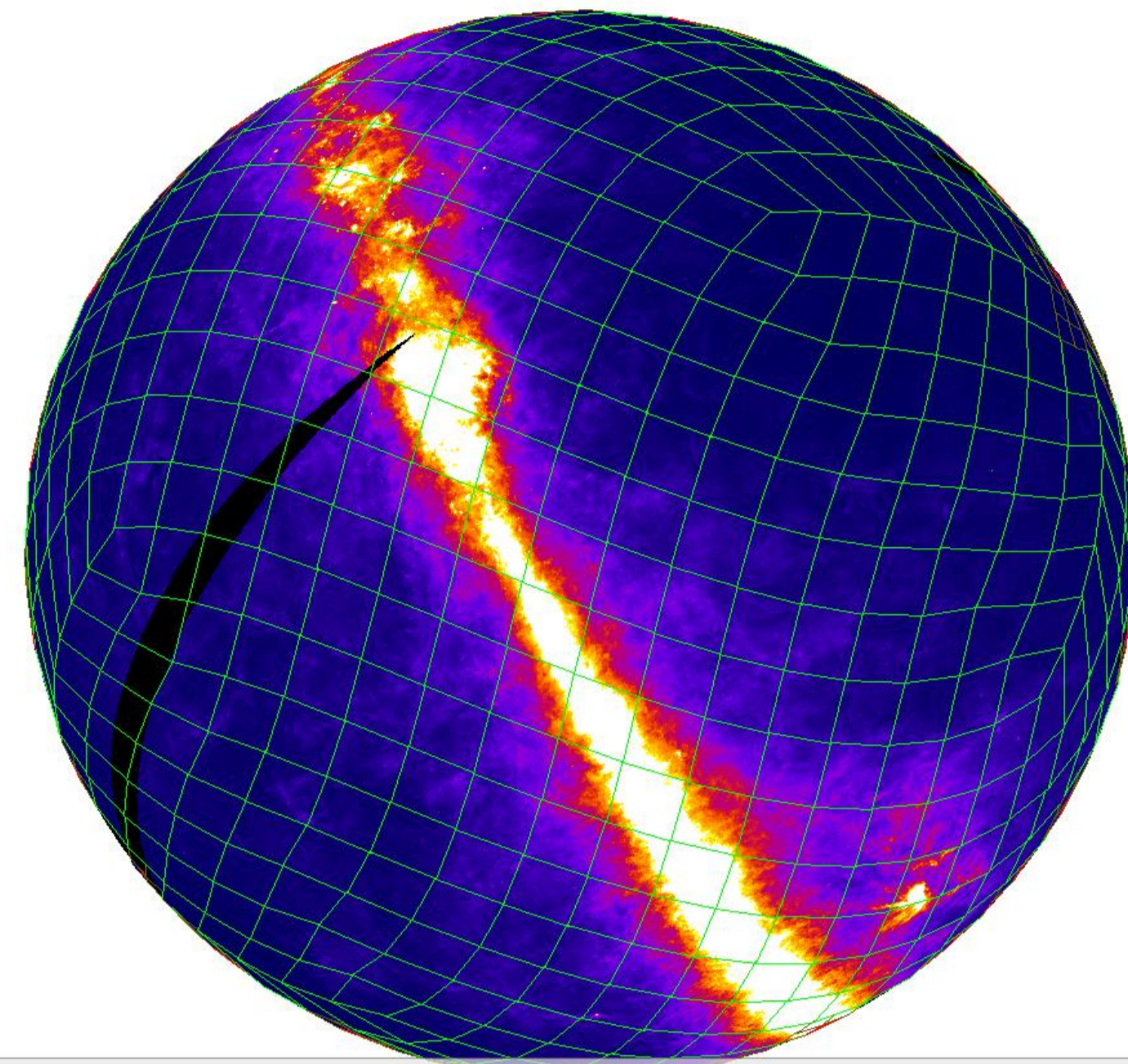


Figure 1: IRIS @100micron HiPS visualisation within Aladin: high resolution on the right

HiPS in two words:

- **HiPS** tiles can be defined for HealPix cells at different resolutions according to Healpix order
- They are hierarchically organized
- Values stored in a given tile are interpolated from original data pixel neighbouring HealPix cell centers values

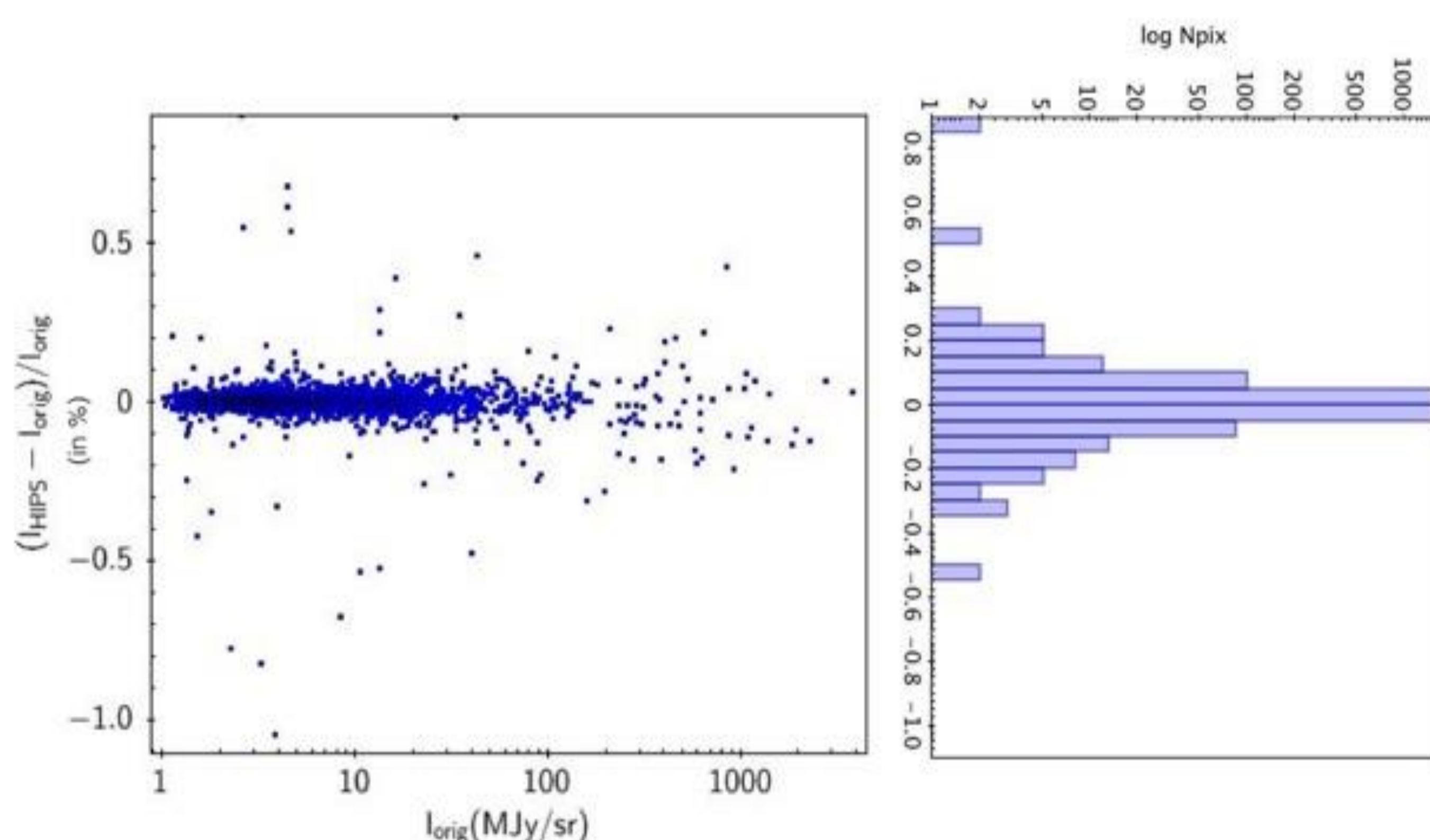


Figure 2: Fractional difference (as a percentage) between the HiPS and original surface brightnesses as a function of the average surface brightness. The histogram of the fractional difference is shown on the right.

Test data and procedures:

- The IRIS data collection is a reprocessing of the IRAS Sky Survey Atlas, recalibrated using DIRBE and made of 12.5 degrees joint images at 12, 25, 60 and 100 μm
- Mean surface brightness is estimated in circular areas on both original images and corresponding areas in the HiPS representation using *Aladin* code controlled via a python module.

Results and conclusions:

- Relative photometric difference is extremely small ($< 0.3\%$!!!) typically much smaller than other sources of uncertainties (fig 2)
- No obvious spatial bias in the results (fig 3) but further tests needed.
- Tests scheduled for source photometry on appropriate surveys
- **For IRAS/IRIS data it is fairly equivalent to measure photometry on data distributed in the HiPS format than on original ones**

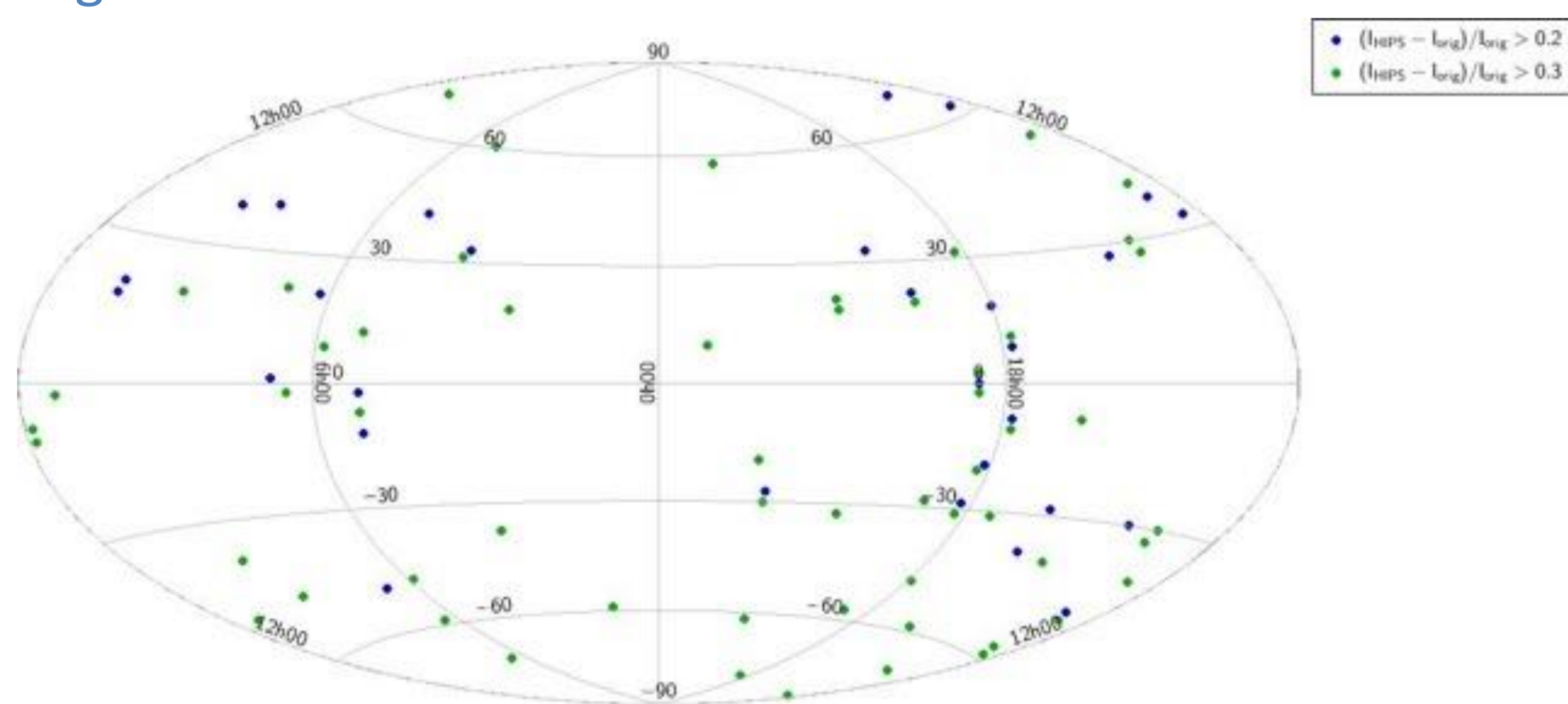


Figure 3: Repartition of higher deviations over the sky